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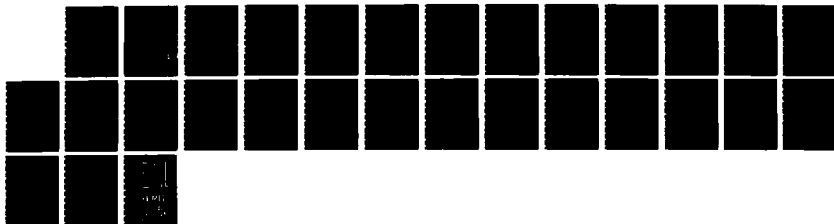
RAPID SOLUTION OF INTEGRAL EQUATIONS OF SCATTERING
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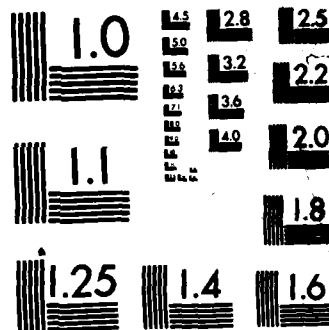
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**Rapid Solution of Integral Equations of
Scattering Theory in Two Dimensions**

V. Rokhlin

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November 1985

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The present paper describes an algorithm for rapid solution of boundary value problems for the Helmholtz equation in two dimensions based on iteratively solving integral equations of scattering theory. CPU time requirements of previously published algorithms of this type are of the order n^3 , where n is the number of nodes in the discretization of the boundary of the scatterer. The CPU time requirements of the algorithm of the present paper are $n^{5/3}$, and can be further reduced, making it considerably more practical for large scale problems.

operators (mathematics),

Keywords: radiation fields

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1. Introduction

One of standard approaches to numerical treatment of boundary value problems for elliptic partial differential equations (PDEs) calls for converting them into second kind integral equations (SKIEs) with subsequent discretization of the latter via appropriate quadrature formulae. Discretization of the resulting SKIEs usually leads to dense large-scale systems of linear algebraic equations, which are in turn solved by means of some iterative technique, such as Generalized Conjugate Residual algorithm (see [11, 21]). Most iterative schemes for the solution of linear systems of this type require application of the matrix of the system to a sequence of recursively generated vectors. Applying a dense matrix to a vector is an order n^2 procedure, where n is the dimension of the matrix, which in this case is equal to the number of nodes in the discretization of the domain of the integral equation. As a result, the whole process is at least of the order n^2 , and for many large scale problems, this estimate is prohibitively large.

In the present paper, we describe an algorithm for rapid application of matrices resulting from discretization of integral equations of scattering theory in two dimensions to arbitrary vectors. The algorithm requires an amount of work proportional to $n^{4/3}$ where n is the number of nodes in the discretization of the boundary of the scatterer, and when it is combined with a Generalized Conjugate Residual type algorithm, the resulting process takes very few iterations to converge, leading to an order $n^{4/3}$ algorithm for the solution of the original scattering problem.

Reduction of boundary value problems for elliptic PDEs to second kind integral equations is discussed in detail in [6, 15, 17]. Numerical treatment of SKIEs in the general case can be found, for example in [3], and numerical solution of acoustic scattering problems in two dimensions by means of SKIEs is discussed in [16]. We present an algorithm for rapid solution of integral equations of classical potential theory (Laplace's equation) in [18], and the algorithm of the present paper can be viewed as an extension of the approach of [18] to the case of the Helmholtz equation. However, the analytical apparatus of the present paper is considerably more complicated than the analytical apparatus of [18], reflecting the difference between the behavior of solutions of the Helmholtz equation and that of harmonic functions.

Remark 1.1

While the algorithm of the present paper has an asymptotic CPU time estimate $n^{4/3}$, it can be easily modified into an order $n \log(n)$ algorithm (see Subsection 4.4). However, it appears that this modification would not lead to significant improvement in actual calculation times for most problems of practicable size ($n \leq 20000$).

2. Background Information

2.1. Notation

We will be considering the situation depicted in Figure 1. A fluid scatterer of arbitrary shape is imbedded in a two-dimensional fluid space. The boundary of the scatterer parametrized by its length will be denoted by γ so that $\gamma : [0, L] \rightarrow \mathbb{R}^2$ is a Jordan curve, and the image of γ will be denoted by Γ . The open interior of γ will be denoted by Ω , so that $\Gamma = \partial\Omega$. We will assume that γ is at least C^2 , i.e., that at each point it has at least two continuous derivatives. The interior normal to γ at the point $x = \gamma(t)$ will be denoted by $N(t)$, and it will always be assumed that $\|N(t)\| = 1$. The density of the scatterer will be denoted by ρ^{in} , and the speed of sound in it will be denoted by c^{in} . The density of the containing space will be denoted by ρ^{out} , and the speed of sound in it will be denoted by c^{out} . We will denote the angular frequency of the source by ω , and its location by x_s . Finally, we will denote the Helmholtz coefficients inside and outside the scatterer by k_{in} and k_{out} respectively (as is well known, $k_{in} = \omega/c^{in}$, $k_{out} = \omega/c^{out}$). In the notation introduced above, we assume that ρ^{in} and ρ^{out} are positive real numbers, and that c^{in} , c^{out} and ω are complex numbers such that $\text{Re}(c^{in}) > 0$, $\text{Re}(c^{out}) > 0$, $\text{Re}(\omega) > 0$ and $\text{Im}(\omega) \geq 0$, $\text{Im}(c^{in}) \leq 0$, $\text{Im}(c^{out}) \leq 0$.

2.2. Single and double layer potentials.

For a Helmholtz equation

$$\nabla^2 \phi + k^2 \phi = 0 \quad (2.1)$$

we will define the field $\phi_{x_0}^k : R^2 \setminus \{x_0\} \rightarrow C^1$ of a unit charge located at the point $x_0 \in R^2$ by the formula

$$\phi_{x_0}^k(x) = H_0(k\|x - x_0\|). \quad (2.2)$$

We will define the field $\phi_{x_0, h}^k$ of a unity dipole located at x_0 and oriented in the direction $h \in R^2$ by the formula

$$\phi_{x_0, h}^k(x) = -H_1(k\|x - x_0\|) \cdot \frac{k(x - x_0, h)}{\|x - x_0\|}. \quad (2.3)$$

For a continuous function $\sigma : [0, L] \rightarrow C^1$, the potential of a single layer of density σ on a curve γ is a mapping $P_{k\sigma}^0 : R^2 \rightarrow C^1$ defined by the formula

$$P_{k\sigma}^0(x) = \int_0^L \phi_{\gamma(t)}^k(x) \sigma(t) dt \quad (2.4)$$

and the potential of a double layer of density σ on a curve γ is a mapping $P_{k\sigma}^1 : R^2 \rightarrow C^1$ defined by the formula

$$P_{k\sigma}^1(x) = \int_0^L \phi_{\gamma(t), N(t)}^k(x) \sigma(t) dt. \quad (2.5)$$

Remark 2.1

Note that while both $P_{k\sigma}^0$ and $P_{k\sigma}^1$ are defined on all of R^2 , neither $P_{k\sigma}^1$ nor the derivatives of $P_{k\sigma}^0$ are continuous in the neighborhood of Γ . The exact nature of their singularities is crucial for the derivation of the equations 2.13, 2.14 of the following subsection, and it is discussed in great detail in [16], [15].

2.3. Acoustic scattering in two dimensions.

In the present paper, we will be considering the following problem:

For a pair of continuous functions $f, g : \Gamma \rightarrow C^1$, find two mappings $\phi : \bar{\Omega} \rightarrow C^1$, $\psi : R^2 \setminus \Omega \rightarrow C^1$ such that

a.

$$\nabla^2 \phi + k_{in}^2 \phi = 0 \quad \text{on } \Omega \quad (2.6)$$

b.

$$\nabla^2 \psi + k_{out}^2 \psi = 0 \quad \text{on } R^2 \setminus \bar{\Omega} \quad (2.7)$$

c.

$$\rho^{in} \cdot \phi + \rho^{out} \cdot \psi = f \quad \text{on } \Gamma \quad (2.8)$$

d.

$$\frac{\partial}{\partial N}(\phi - \psi) = g \quad \text{on } \Gamma \quad (2.9)$$

e. ψ satisfies the radiation condition at ∞ , i.e., for any $x \in R^2$, there exists $c \in C^1$ such that

$$\lim_{t \rightarrow \infty} \psi(t \cdot x) \cdot e^{-ikt \cdot |x|} \cdot \sqrt{t} = c \quad (2.10)$$

with $k = k_{out}$.

The above five equations describe acoustic scattering from a two-dimensional fluid inclusion in a fluid space in the frequency domain, and have been studied in great detail (see, for example, [2],

[5], [10], [16]. Their numerical solution, however, presents a number of serious difficulties, especially for large-scale problems. Here we will follow the approach of [16], which calls for reducing these equations to second kind integral equations and solving the latter numerically. As is shown in [16], by introducing two new unknown functions $\sigma, \eta : [0, L] \rightarrow C^1$ and representing the functions ϕ, ψ in the form

$$\phi = \frac{1}{\rho^{in}} \cdot P_{kin}^0 \eta + P_{kin}^1 \sigma \quad (2.11)$$

$$\psi = \frac{1}{\rho^{out}} \cdot P_{kout}^0 \eta + P_{kout}^1 \sigma \quad (2.12)$$

the equations (2.1)–(2.5) are reduced to a pair of second kind integral equations on the boundary of the scatterer:

$$-2i(\rho^{out} + \rho^{in})\sigma + (P_{kout}^0 \eta - P_{kin}^0 \eta) + (\rho^{out} P_{kout}^1 \sigma - \rho^{in} P_{kin}^1 \sigma) = f, \quad (2.13)$$

$$2i\left(\frac{1}{\rho^{out}} + \frac{1}{\rho^{in}}\right)\eta + \frac{\partial}{\partial N}(P_{kout}^1 \sigma P_{kout}^1 \sigma) + \frac{\partial}{\partial N}\left(\frac{1}{\rho^{out}} P_{kout}^0 \eta - \frac{1}{\rho^{in}} P_{kin}^0 \eta\right) = g. \quad (2.14)$$

2.4. Iterative solution of second kind integral equations.

The system of equations (2.13), (2.14) satisfies the conditions of the Fredholm theorems and can be efficiently solved by means of Generalized Conjugate Residual type iterative algorithms (see [11], [16], [21]). Iterative solution of integral equations usually involves application of the integral operator in the left-hand side of the equation to a sequence of recursively generated functions. Applying an integral operator to a function numerically is, generally speaking, an order n^2 procedure, where n is the number of nodes in the discretization of the domain of the operator. The resulting CPU time estimate for the solution of the original scattering problem is also of the order n^2 (see [3] [16]), which can be prohibitively expensive for large-scale problems. The rest of this paper is devoted to constructing an algorithm for numerically applying the integral operators in the left-hand side of the equations (2.13), (2.14) to arbitrary functions in a "fast" manner, i.e. for a cost less than n^2 (the particular algorithm we have tested has an asymptotic CPU time estimate $n^{4/3}$).

Remark 2.2

Evaluating integral operators in the left-hand sides of equations (2.13), (2.14) numerically can be viewed as evaluating the fields and normal derivatives of the fields created on the curve γ by charge and dipole distributions on that same curve. In the following section, we develop an analytical apparatus for rapid evaluation of fields (and derivatives of the fields) of distributions of charges and dipoles, and in Section IV, we use this apparatus to construct an actual algorithm for rapid evaluation of integral operators of the forms (2.13), (2.14).

2.5. Asymptotic behavior of Bessel functions.

In agreement with the standard practice, we will denote by J_m the Bessel function of the first kind of order m , and by H_m , the Hankel function of order m . As is well known (see, for example [20]), J_m are analytic on the whole complex plane for all values of m , while H_m have a branch cut along the negative real axis, and become infinite at the origin. The asymptotic behaviour of the functions J_m, H_m for large m is given by the formulae

$$\lim_{m \rightarrow \infty} J_m(z) \cdot \left(\frac{2m}{ez}\right)^m \cdot \sqrt{(2\pi m)} = 1 \quad (2.15)$$

and

$$\lim_{m \rightarrow \infty} H_m(z) \cdot \left(\frac{ez}{2m}\right)^m \cdot \frac{\sqrt{(\pi m)}}{\sqrt{2}} = -1 \quad (2.16)$$

(see [1], 9.3.1, 9.3.2, 9.1.3). For large z and fixed m , the asymptotic behavior of $J_m(z)$, $H_m(z)$ is given by the formulae

$$\sqrt{z}J_m(z) - \sqrt{\frac{2}{\pi}} \cos\left(z - \frac{m\pi}{2} - \frac{\pi}{4}\right) = O\left(\frac{e^{Im(z)}}{|z|}\right), \quad (2.17)$$

$$\sqrt{z}H_m(z) - \sqrt{\frac{2}{\pi}} e^{i(z - \frac{m\pi}{2} - \frac{\pi}{4})} = O\left(\frac{e^{-Im(z)}}{|z|}\right) \quad (2.18)$$

when $z \rightarrow \infty$, as long as $Im(z) \geq 0$ (see [1], 9.2.5, 9.2.7).

3. Rapid Evaluation of Radiation Fields.

3.1. Partial wave expansions of radiation fields.

If a function $\phi : R^2 \rightarrow C^1$ satisfies the Helmholtz equation (2.1) in an open disk D of radius R with the center at the point $x_0 \in R^2$ then there exists a unique sequence $\alpha = \{\alpha_m\}$, $m = 0, \pm 1, \pm 2, \dots$ such that for any $x \in D$,

$$\phi(x) = \sum_{m=-\infty}^{+\infty} \alpha_m J_m(k\rho) e^{im\theta}. \quad (3.1)$$

In the above formula, $\rho = \|x - x_0\|$ and θ is the angle between the vector $x - x_0$ and the x axis.

If a function ψ satisfies the equation (2.1) outside D and the radiation condition (2.10) at ∞ then there exists a unique sequence $\beta = \{\beta_m\}$, $m = 0, \pm 1, \pm 2, \dots$ such that for any $x \in R^2 \setminus \bar{D}$,

$$\psi(x) = \sum_{m=-\infty}^{+\infty} \beta_m H_m(k\rho) e^{im\theta}. \quad (3.2)$$

A derivation of the formulae (3.1), (3.2) can be found, for example, in [15], and we will refer to functions satisfying the Helmholtz equation as radiation fields, to expansions of forms (3.1), (3.2) as J-expansions and H-expansions respectively, and to the point x_0 as the center of the expansions (3.1), (3.2).

The following lemma is a direct consequence of the formulae (2.15), (2.16). It establishes the convergence rates of the expansions (3.1), (3.2).

Lemma 3.1.

If $D_1 \subset D$ is a disk of radius $R_1 < R$ with the center at x_0 then there exists $c > 0$ such that for any $x \in D_1$ and $N > |k| \cdot R_1$,

$$|\phi(x) - \sum_{m=-N}^N \alpha_m J_m(k\rho) e^{im\theta}| < c \left(\frac{R_1}{R}\right)^N. \quad (3.3)$$

If $D_2 \supset D$ is a disk of radius $R_2 > R$ with the center at x_0 then there exists $c > 0$ such that for any $x \in R^2 \setminus \bar{D}_2$ and $N > |k| \cdot R$,

$$|\psi(x) - \sum_{m=-N}^N \beta_m H_m(k\rho) e^{im\theta}| < c \left(\frac{R}{R_2}\right)^N. \quad (3.4)$$

Remark 3.1

In numerical calculations, expansions (3.1), (3.2) are truncated after a finite number of terms, and the resulting expressions are viewed as approximations to the fields ϕ , ψ . If we want to approximate ϕ by an expansion of the form (3.3) with an accuracy ϵ then according to the above lemma, we have to choose

$$N \geq \max(|k| \cdot R_1, \frac{-\ln(\epsilon) + \ln(c)}{\ln(R) - \ln(R_1)}). \quad (3.5)$$

Since logarithm is a very slowly growing function, for medium and large scale problems,

$$\max(R_1 \cdot |k|, \frac{-\ln(\epsilon) + \ln(c)}{\ln(R) - \ln(R_1)}) \sim R_1 \cdot |k|.$$

i.e. the number of terms in the approximation is almost independent of ϵ , and must be roughly equal to $|k|R_1$. A similar calculation shows that for medium to large scale problems, the expansion (3.2) can be truncated after approximately $N = |k|R$ terms.

3.2. Translation operators for H and J-expansions.

For a real number $r > 0$, we will denote by X_r the linear space of all two-sided complex sequences $\alpha = \{\alpha_m\}$, $m = 0, \pm 1, \pm 2, \dots$ such that for some $c > 0$,

$$|\alpha_m| \cdot \left(\frac{2m}{er}\right)^m \cdot \sqrt{m} < c \quad (3.6)$$

for all $m \geq r$. We will denote by Y_r the linear space of all complex sequences $\beta = \{\beta_m\}$, $m = 0, \pm 1, \pm 2, \dots$, such that for some $c > 0$,

$$|\beta_m| \cdot \left(\frac{er}{2m}\right)^m \cdot \sqrt{m} < c \quad (3.7)$$

for all $m \geq r$. It is easy to see that $X_r \subset Y_r$, and that the condition (3.6) is a very restrictive one, since in order to satisfy it, the elements of the sequence $\{\alpha_m\}$ must decay roughly as $(r/2)^m \cdot m!$, while the condition (3.7) is a very mild one - it prohibits the elements of $\{\beta_m\}$ from growing faster than approximately $(2/r)^m \cdot m!$. By applying formulae (9.3.1), (9.3.2) from [1], it is easy to show that in (3.1), (3.2), $\alpha \in Y_{|k|R}$ and $\beta \in Y_{|k|R}$. Conversely, for any sequence $\alpha \in Y_{|k|R}$, the expansion (3.1) converges inside D , and for any $\beta \in Y_{|k|R}$, the expansion (3.2) converges outside D . For a natural n , we will denote by T_n a linear mapping $Y_r \rightarrow Y_r$ converting a sequence $\alpha = \{\alpha_m\}$, $m = 0, \pm 1, \pm 2, \dots$ into a sequence $\tilde{\alpha} = \{\tilde{\alpha}_m\}$, $m = 0, \pm 1, \pm 2, \dots$ defined by the formulae

$$\begin{aligned} \tilde{\alpha}_m &= \alpha_m & \text{for } |m| \leq n \\ \tilde{\alpha}_m &= 0 & \text{for } |m| \geq n+1. \end{aligned}$$

Clearly, $T_n(Y_r) \subset X_r$, and for obvious reasons, we will refer to T_n as truncation.

For the remainder of this section, D_1, D_2, D_3 will denote three disks in R^2 such that $D_2 \subset D_1$ and $D_1 \cap D_3 = \emptyset$. The centers and radii of these disks will be denoted by c_1, c_2, c_3 and R_1, R_2, R_3 respectively. We will denote the distance $\|c_2 - c_1\|$ by ρ_{12} , and the distance $\|c_3 - c_1\|$ by ρ_{13} . The angle between the vector $c_2 - c_1$ and the x axis will be denoted by θ_{12} , and the angle between $c_3 - c_1$ and the x axis will be denoted by θ_{13} . For a point $x \in R^2$, we will denote $\|x - c_1\|$ by ρ_1 , $\|x - c_2\|$ by ρ_2 , and $\|x - c_3\|$ by ρ_3 . Finally, the angles between the vectors $x - c_1, x - c_2$ and $x - c_3$ and the x axis will be denoted by $\theta_1, \theta_2, \theta_3$ respectively.

Suppose that n is a natural number or ∞ . We will define a linear operator $U_{c_2 c_1}^n : X_{|k|R_2} \rightarrow X_{|k|R_1}$ as follows. If $A = \{a_m\}$, $m = 0, \pm 1, \pm 2, \dots$ is an element of $X_{|k|R_2}$ then $U_{c_2 c_1}^n(A) = B = \{b_m\}$, $m = 0, \pm 1, \pm 2, \dots$ is defined by the formula

$$b_m = \sum_{j=-n}^n e^{-ij\theta_{12}} a_{m-j} J_j(k\rho_{12}) \quad (3.8)$$

for all $m = 0, \pm 1, \pm 2, \dots$. An operator $V_{c_1 c_2}^n : Y_{|k|_{R_1}} \rightarrow Y_{|k|_{R_2}}$ will be defined by the formula

$$b_m = \sum_{j=-n}^n e^{-ij(\theta_{12}-\pi)} a_{m-j} J_j(k\rho_{12}) \quad (3.9)$$

for any $A = \{a_m\}$, $B = \{b_m\}$ such that $B = V_{c_1 c_2}^n(A)$. Finally, the operator $X_{|k|_{R_1}} \rightarrow Y_{|k|_{R_2}}$ converting an arbitrary sequence $A = \{a_m\} \in X_{|k|_{R_1}}$ into the sequence $B = \{b_m\} \in Y_{|k|_{R_2}}$ defined by the formula

$$b_m = \sum_{j=-n}^n e^{-ij(\theta_{12}-\pi)} a_{m-j} H_j(k\rho_{13}) \quad (3.10)$$

will be denoted by $W_{c_1 c_3}^n$. For any natural m , we will define the operators $U_{c_2 c_1}^{nm} : X_{|k|_{R_2}} \rightarrow X_{|k|_{R_1}}$, $V_{c_1 c_2}^{nm} : Y_{|k|_{R_1}} \rightarrow Y_{|k|_{R_2}}$, $W_{c_1 c_3}^{nm} : X_{|k|_{R_1}} \rightarrow Y_{|k|_{R_2}}$ by the formulae

$$U_{c_2 c_1}^{nm} = T_m \cdot U_{c_2 c_1}^n \cdot T_m,$$

$$V_{c_1 c_2}^{nm} = T_m \cdot V_{c_1 c_2}^n \cdot T_m,$$

$$W_{c_1 c_3}^{nm} = T_m \cdot W_{c_1 c_3}^n \cdot T_m.$$

Remark 3.2

We will denote by Y the set of all two-sided complex sequences $\{y_i\}$, $i = \pm 1, \pm 2, \dots$. When n is a natural number (as opposed to ∞), the formulae (3.8), (3.9), (3.10) define mappings $\tilde{U}_{c_2 c_1}^n, \tilde{V}_{c_1 c_2}^n, \tilde{W}_{c_1 c_3}^n : Y \rightarrow Y$ that can be viewed as extensions of the mappings $U_{c_2 c_1}^n, V_{c_1 c_2}^n, W_{c_1 c_3}^n$ i.e.,

$$\tilde{U}_{c_2 c_1|X_{|k|_{R_2}}}^n = U_{c_2 c_1}^n, \quad \tilde{V}_{c_1 c_2|X_{|k|_{R_1}}}^n = V_{c_1 c_2}^n, \quad \tilde{W}_{c_1 c_3|X_{|k|_{R_1}}}^n = W_{c_1 c_3}^n.$$

Whenever there is no danger of confusion, we will make no distinction between the mappings $\tilde{U}_{c_2 c_1}^n, \tilde{V}_{c_1 c_2}^n, \tilde{W}_{c_1 c_3}^n$ and the mappings $U_{c_2 c_1}^n, V_{c_1 c_2}^n, W_{c_1 c_3}^n$ respectively.

The following three lemmas justify our referring to the operators $U_{c_2 c_1}^n, V_{c_1 c_2}^n, W_{c_1 c_3}^n, U_{c_2 c_1}^{nm}, V_{c_1 c_2}^{nm}, W_{c_1 c_3}^{nm}$ as translation operators, and will be used to shift the origins of H and J -expansions, and to convert H -expansions into J -expansions. They are a direct consequence of the Graf's addition theorem for Bessel functions (see [1], (9.1.79)).

Lemma 3.2. Suppose that $\psi : R^2 \rightarrow C^1$ is a radiation field analytical in $R^2 \setminus \bar{D}_2$ and satisfying the radiation condition (2.10) at ∞ . Suppose further that ψ is represented by an expansion

$$\psi(x) = \sum_{m=-\infty}^{+\infty} \beta_m H_m(k\rho_2) e^{im\theta_2} \quad (3.11)$$

valid in $R^2 \setminus \bar{D}_2$, and by an expansion

$$\psi(x) = \sum_{m=-\infty}^{+\infty} \tilde{\beta}_m H_m(k\rho_1) e^{im\theta_1} \quad (3.12)$$

valid in $R^2 \setminus \bar{D}_1$. Then $\{\tilde{\beta}_m\} = U_{c_2 c_1}^\infty(\{\beta_m\})$.

Lemma 3.3. Suppose that $\psi : R^2 \rightarrow C^1$ is a radiation field analytical in D_1 and represented by an expansion

$$\phi(x) = \sum_{m=-\infty}^{+\infty} \alpha_m J_m(k\rho_1) e^{im\theta_1} \quad (3.13)$$

valid in D_1 , and by an expansion

$$\phi(x) = \sum_{m=-\infty}^{+\infty} \tilde{\alpha}_m J_m(k\rho_2) e^{im\theta_2} \quad (3.14)$$

valid in D_2 . Then $\{\tilde{\alpha}_m\} = V_{c_1 c_2}^\infty(\{\alpha_m\})$.

Lemma 3.4. Suppose that $\psi : R^2 \rightarrow C^1$ is a radiation field analytical outside the disk D_1 and satisfying the radiation condition (2.10), and that it is represented in $R^2 \setminus \bar{D}_1$ by the expansion (3.12). Then inside the disk D_3 , the function ψ can be represented in the form

$$\psi(x) = \sum_{m=-\infty}^{+\infty} \gamma_m J_m(k\rho_3) e^{im\theta_3} \quad (3.15)$$

with $\{\gamma_m\} = W_{c_1 c_3}^\infty(\{\tilde{\beta}_m\})$.

Under the conditions of Lemma 3.2, we will define a radiation field $\phi_{c_2 c_1}^n : R^2 \setminus \bar{D}_1 \rightarrow C^1$ by the expression

$$\psi_{c_2 c_1}^n(x) = \sum_{m=-\infty}^{+\infty} \delta_m H_m(k\rho_1) e^{im\theta_1} \quad (3.16)$$

with the coefficients $\{\delta_m\}$, $m = 0, \pm 1, \pm 2, \dots$ defined by the formula $\{\delta_m\} = U_{c_2 c_1}^n(\{\beta_m\})$. Similarly, under the conditions of Lemma 3.3, we will define a radiation field $\psi_{c_1 c_2}^n : D_2 \rightarrow C^1$ by the formula

$$\phi_{c_1 c_2}^n(x) = \sum_{m=-\infty}^{+\infty} \tilde{\delta}_m J_m(k\rho_2) e^{im\theta_2} \quad (3.17)$$

with $\{\tilde{\delta}_m\} = V_{c_1 c_2}^n(\{\alpha_m\})$. Finally, under the conditions of Lemma 3.4, we will denote by $\bar{\psi}_{c_1 c_3}^n$ the radiation field $D_3 \rightarrow C^1$ defined by the formula

$$\bar{\psi}_{c_1 c_3}^n(x) = \sum_{m=-\infty}^{+\infty} \bar{\delta}_m J_m(k\rho_3) e^{im\theta_3} \quad (3.18)$$

with the sequence $\{\bar{\delta}_m\}$, $m = 0, \pm 1, \pm 2, \dots$ defined by the formula $\{\bar{\delta}_m\} = W_{c_1 c_3}^n(\{\beta_m\})$.

Obviously, $\psi_{c_2 c_1}^\infty(x) = \psi(x)$ for any $x \in R^2 \setminus \bar{D}_1$, $\phi_{c_2 c_1}^\infty(x) = \phi(x)$ for any $x \in D_2$, and $\bar{\psi}_{c_1 c_3}^\infty(x) = \psi(x)$ for any $x \in D_3$, and we will view the mappings $\psi_{c_2 c_1}^n$, $\phi_{c_1 c_2}^n$, $\bar{\psi}_{c_1 c_3}^n$ as approximations to the mappings ψ , ϕ and ψ respectively.

3.3. Diagonalization of translation operators.

Theorem 3.1. For any natural n , each of the operators $U_{c_2 c_1}^n$, $V_{c_1 c_2}^n$, $W_{c_1 c_3}^n$, is a bounded scalar type operator (see [8], v. III). Their sets of eigenvectors coincide, and are given by the formula

$$e_q = \{e^{iqm}\}, \quad m = 0, \pm 1, \pm 2, \dots \quad (3.19)$$

with $q \in [0, 2\pi]$. The spectral representations of the operators $U_{c_2 c_1}^n$, $V_{c_1 c_2}^n$, $W_{c_1 c_3}^n$ are respectively

$$U_{c_2 c_1}^n = \int_0^{2\pi} \lambda_n(q) P_q dq, \quad (3.20)$$

$$V_{c_1 c_2}^n = \int_0^{2\pi} \mu_n(q) P_q dq, \quad (3.21)$$

$$W_{c_1 c_3}^n = \int_0^{2\pi} \nu_n(q) P_q dq, \quad (3.22)$$

where P_q is a rank one operator projecting Y on its subspace spanned by the vector e_q and commuting with each of the operators $U_{c_2 c_1}^n$, $V_{c_1 c_2}^n$, $W_{c_1 c_3}^n$, and the functions $\lambda_n, \mu_n, \nu_n : [0, 2\pi] \rightarrow C^1$ are defined by the formulae

$$\lambda_n(q) = \sum_{m=-n}^n e^{-im(q+\theta_{12})} J_m(k\rho_{12}), \quad (3.23)$$

$$\mu_n(q) = \sum_{m=-n}^n e^{-im(q+\theta_{12}-\pi)} J_m(k\rho_{12}), \quad (3.24)$$

$$\nu_n(q) = \sum_{m=-n}^n e^{-im(q+\theta_{13}-\pi)} H_m(k\rho_{13}). \quad (3.25)$$

Proof. An inspection of the formulae (3.8), (3.9), (3.10) shows that the mappings $U_{c_2 c_1}^n$, $V_{c_1 c_2}^n$, $W_{c_1 c_3}^n$ are convolutions of the sequence $\{a_m e\}$, $m = 0, \pm 1, \pm 2, \dots$ with the finite sequences

$$\{e^{-im\theta_{12}} J_m(k\rho_{12})\}, \quad m = 0, \pm 1, \pm 2, \dots \quad (3.26)$$

$$\{e^{-im(\theta_{12}-\pi)} J_m(k\rho_{12})\}, \quad m = 0, \pm 1, \pm 2, \dots \quad (3.27)$$

$$\{e^{-im(\theta_{13}-\pi)} H_m(k\rho_{13})\}, \quad m = 0, \pm 1, \pm 2, \dots \quad (3.28)$$

respectively. Convolutions with finite sequences are diagonalized by the Fourier Transformation (see, for example, [8, 14]), which proves a). We obtain b), c), and d) by applying the Fourier Transformation to the sequences (3.26), (3.27), (3.28).

Remark 3.3

For the mappings $U_{c_2 c_1}^n$, $V_{c_1 c_2}^n$, the above theorem can be extended to the case $n = \infty$. As is well known, for any $z \in C^1$, $\theta \in [0, 2\pi]$,

$$\sum_{m=-\infty}^{+\infty} J_m(z) e^{im\theta} = e^{iz \cos \theta} \quad (3.29)$$

(see, for example, [1], 9.1.21), and for $n = \infty$ the expressions (3.20), (3.21) assume the form

$$\lambda_\infty(q) = e^{ik\rho_{12} \cos(q+\theta_{12})}, \quad (3.30)$$

$$\mu_\infty(q) = e^{ik\rho_{12} \cos(q+\theta_{12}-\pi)}. \quad (3.31)$$

However, Theorem 3.1 can not be extended to the case of the operator $W_{c_1 c_3}^\infty$ since the latter is unbounded (see [8], VIII).

Remark 3.4

In numerical calculations, the operators $U_{c_2c_1}^{nm}, V_{c_1c_2}^{nm}, W_{c_1c_3}^{nm}$, with sufficiently large n, m will be viewed as approximations to the operators $U_{c_2c_1}^n, V_{c_1c_2}^n, W_{c_1c_3}^n$ (see Remark 3.1). Clearly, applying either of the operators $U_{c_2c_1}^{nm}, V_{c_1c_2}^{nm}, W_{c_1c_3}^{nm}$ numerically to an arbitrary sequence is an order nm procedure, which can be prohibitively expensive for large scale problems. However, it follows from the above theorem that the operators $U_{c_2c_1}^{nm}, V_{c_1c_2}^{nm}, W_{c_1c_3}^{nm}$ are convolutions of sequences of lengths n, m , and such convolutions can be evaluated in order $(n+m) \log(n+m)$ operations by means of the Fast Fourier Transformation (see [4, 12]).

3.4. Asymptotic forms of radiation fields.

In this subsection, we introduce an alternative form of the expansions (3.1), (3.2) and the mappings $U_{c_2c_1}^n, V_{c_1c_2}^n, W_{c_1c_3}^n$, providing a simple physical interpretation of the Lemmas 3.2 - 3.4 and the Theorem 3.1.

For the expansion (3.2), we will consider a function $F_{\psi, x_0} : [0, 2\pi] \rightarrow C^1$ defined by the formula

$$F_{\psi, x_0}(\theta) = \lim_{t \rightarrow \infty} \psi(t \cdot x + x_0) \cdot \sqrt{t} \cdot e^{-ikt} \cdot \frac{\sqrt{k\pi}}{\sqrt{2}} \cdot e^{\frac{\pi}{4}i} \quad (3.32)$$

with $x = (\cos \theta, \sin \theta)$. Substituting (2.18), (3.2), into (3.32), we obtain

$$F_{\psi, x_0}(\theta) = \sum_{m=-\infty}^{+\infty} \beta_m e^{-\frac{m\pi}{2}i} e^{im\theta} \quad (3.33)$$

which provides an explicit expression of F_{ψ, x_0} via the coefficients $\{\beta_m\}$, and we will refer to F_{ψ, x_0} as the asymptotic representation of the field ψ with the origin at x_0 .

In order to define an asymptotic representation of the field ϕ in (3.1), we will have to introduce an additional assumption that

$$\sum_{m=-\infty}^{+\infty} |\alpha_m| = c < \infty \quad (3.34)$$

which guarantees that the expansion (3.1) converges on the whole R^2 . By combining (2.17) and (3.1), it is easy to show that for any $\theta \in [0, 2\pi]$ and $x = (\cos \theta, \sin \theta)$, there exist unique numbers U_θ, V_θ such that

$$\lim_{t \rightarrow \infty} \psi(t \cdot x + x_0) \cdot \sqrt{(2k\pi t)} - (e^{(kt - \frac{\pi}{4})i} \cdot U_\theta + e^{-(kt - \frac{\pi}{4})i} \cdot V_\theta \cdot e^{ikt} = 0), \quad (3.35)$$

and that the numbers U_θ, V_θ are defined by the formulae

$$U_\theta = \sum_{m=-\infty}^{+\infty} (\alpha_m \cdot e^{-\frac{m\pi}{2}i}) \cdot e^{im\theta}, \quad (3.36)$$

$$V_\theta = \sum_{m=-\infty}^{+\infty} (\alpha_m \cdot e^{\frac{m\pi}{2}i}) \cdot e^{im\theta}. \quad (3.37)$$

Now, we will define the function $G_{\phi, x_0} : [0, 2\pi] \rightarrow C^1$ by the formula

$$G_{\phi, x_0}(\theta) = U_\theta \quad (3.38)$$

and refer to it as the asymptotic representation of the field ψ with the origin at x_0 .

The following three lemmas establish that the asymptotic representations of the fields ψ, ϕ diagonalize the translation operators $U_{c_2c_1}^n, V_{c_1c_2}^n, W_{c_1c_3}^n$. All three are an immediate consequence of the Lemmas 3.2-3.4 and the Theorem 3.1.

Lemma 3.5.

Under the conditions of Lemma 3.2, for any $\theta \in [0, 2\pi]$,

$$F_{\psi, c_1}(\theta) = F_{\psi, c_2}(\theta) \cdot \lambda_{\infty}(\theta) \quad (3.39)$$

with $\lambda_{\infty} : [0, 2\pi] \rightarrow C^1$ defined by (3.30).

Lemma 3.6. Under the conditions of Lemma 3.3, for any $\theta \in [0, 2\pi]$,

$$G_{\psi, c_2}(\theta) = G_{\psi, c_1}(\theta) \cdot \mu_{\infty}(\theta) \quad (3.40)$$

with $\mu_{\infty} : [0, 2\pi] \rightarrow C^1$ defined by (3.31).

Lemma 3.7. Under the conditions of Lemma 3.4, for any $\epsilon > 0$ there exists $N > 0$ such that for any $n \geq N$,

$$|\psi(x) - \phi_n(x)| < \epsilon \quad (3.41)$$

for any $x \in D_3$ where $\phi_n : R^2 \rightarrow C^1$ is a radiation field analytical inside D_3 with the asymptotic representation defined by the formula

$$G_{\psi_n, c_3} = \nu_n(\theta) F_{\psi, c_1}(\theta), \quad (3.42)$$

and the function $\nu_n : [0, 2\pi] \rightarrow C^1$ is defined by (3.22).

The following two lemmas are a direct consequence of the Theorem 3.1, Lemmas 3.5, 3.6 and the Remark 3.3. They provide explicit expressions for the asymptotic representations of the fields generated by a charge and a dipole located at an arbitrary point, and for the values and the derivatives of a field given by its asymptotic representation.

Lemma 3.8. Suppose that under the conditions of the Lemma 3.2, the field ψ is defined by the formula $\psi = \phi_x^k$, i.e., it is the field of a unity charge located at the point x . Then the asymptotic representation of the field ψ with the center c_1 is given by the formula

$$F_{\psi, c_1}(\theta) = e^{-ik\rho_1 \cos(\theta - \theta_1)}. \quad (3.43)$$

If the field ψ is given by the formula $\psi = \phi_{x, h}^k$, i.e., it is the field of a unity dipole located at the point x and oriented in the direction h , then the asymptotic representation of (3.12) with the center at c_1 is given by the formula

$$F_{\psi, c_1}(\theta) = ik \cos(\theta - \theta_h) e^{-ik\rho_1 \cos(\theta - \theta_1)} \quad (3.44)$$

where θ_h is the angle between the vector h and the x axis.

Lemma 3.9. Suppose that under the conditions of the Lemma 3.3, the field (3.13) has the asymptotic representation $G_{\psi, c_1} : [0, 2\pi] \rightarrow C^1$. Then for any $x \in D_1$, $h \in R^2$ such that $\|h\| = 1$,

$$\phi(x) = \frac{1}{2\pi} \int_0^{2\pi} G_{\psi, c_1}(\theta) \cdot e^{-ik\rho_1 \cos(\theta - \theta_1)} d\theta, \quad (3.45)$$

$$\frac{d}{dt} \phi(x + th)|_{t=0} = \frac{ik}{2\pi} \int_0^{2\pi} G_{\psi, c_1}(\theta) \cos(\theta - \theta_h) e^{-ik\rho_1 \cos(\theta - \theta_1)} d\theta. \quad (3.46)$$

3.5. Numerical evaluation of translation operators.

For the rest of this paper, we will view the asymptotic representations (3.32), (3.38) (as opposed to the expansions of the forms (3.1), (3.2)) as our principal tool for representing radiation fields. Lemma 3.8 permits one to calculate asymptotic representations of fields of distributions of charges and dipoles without evaluating the coefficients of their H -expansions, and Lemma 3.9 provides a tool for calculating the fields and derivatives of the fields with given asymptotic representations without having to evaluate the coefficients of J -expansions of these fields.

For a radiation field $\psi : R^2 \rightarrow C^1$ analytical outside D_1 and satisfying the radiation condition (2.10), and an integer $n \geq 2$, we will denote by F_{ψ, c_1}^n the finite sequence a_1, a_2, \dots, a_n defined by the formulae

$$a_i = F_{\psi, c_1}(w_i), \quad (3.47)$$

$$w_i = 2\pi \frac{(i-1)}{n}, \quad i = 1, 2, \dots, n. \quad (3.48)$$

Similarly, for a radiation field ϕ analytical inside D_1 and possessing an asymptotic representation G_{ϕ, c_1} we will denote by G_{ϕ, c_1}^n the sequence b_1, b_2, \dots, b_n defined by the formula

$$b_i = G_{\phi, c_1}(w_i). \quad (3.49)$$

Obviously, $F_{\psi, c_1}^n, G_{\phi, c_1}^n$ are tabulations at n equispaced nodes on the interval $[0, 2\pi]$ of the functions $F_{\psi, c_1}, G_{\phi, c_1}$ respectively, and we will view $F_{\psi, c_1}^n, G_{\phi, c_1}^n$ as finite-dimensional projections of the asymptotic representations of the ψ, ϕ radiation fields.

For a finite sequence $G_{\phi, c_1}^n = \{b_1, b_2, \dots, b_n\}$ we will consider a radiation field $\tilde{G}_{\psi, c_1}^n : R^2 \rightarrow C^1$ defined by the formula

$$\tilde{G}_{\phi, c_1}^n(x) = \frac{1}{n} \sum_{j=1}^n b_j e^{-ik\rho_1 \cos(w_j - \theta_1)}. \quad (3.50)$$

Clearly, (3.50) is a trapezoidal approximation to the integral (3.45), and we will look upon (3.50) as an approximation to the field ϕ . Differentiating (3.50) with respect to x , we obtain the formula

$$\frac{d}{dt} \tilde{G}_{\phi, c_1}^n(x + th)|_{t=0} = \frac{ik\|h\|}{n} \sum_{j=1}^n b_j \cos(w_j - \theta_h) e^{-ik\rho_1 \cos(w_j - \theta_1)} \quad (3.51)$$

for any $h \in R^2$, and we will view (3.51) as an approximation to (3.46).

Finally, we will define mappings $P_{c_2 c_1}^{mn}, Q_{c_1 c_2}^{mn}, S_{c_1 c_3}^{mn} : C^n \rightarrow C^n$ by the formulae

$$P_{c_2 c_1}^{mn}(z_1, z_2, \dots, z_n) = (\lambda_m(w_1) \cdot z_1, \lambda_m(w_2) \cdot z_2, \dots, \lambda_m(w_n) \cdot z_n), \quad (3.52)$$

$$Q_{c_1 c_2}^{mn}(z_1, z_2, \dots, z_n) = (\mu_m(w_1) \cdot z_1, \mu_m(w_2) \cdot z_2, \dots, \mu_m(w_n) \cdot z_n), \quad (3.53)$$

$$S_{c_1 c_3}^{mn}(z_1, z_2, \dots, z_n) = (\nu_m(w_1) \cdot z_1, \nu_m(w_2) \cdot z_2, \dots, \nu_m(w_n) \cdot z_n), \quad (3.54)$$

with the functions λ_m, μ_m, ν_m defined by (3.20), (3.21), (3.22). It is easy to see from the formulae (3.22)–(3.23), (3.52)–(3.54) and the Theorem 3.1 that under the conditions of the Lemmas 3.2–3.4,

$$P_{c_2 c_1}^{mn}(F_{\psi, c_2}^n) = F_{U, c_1}^n,$$

$$Q_{c_1 c_2}^{mn}(G_{\psi, c_1}^n) = G_{V, c_2}^n,$$

$$S_{c_1 c_3}^{mn}(F_{\psi, c_1}^n) = G_{W, c_3}^n,$$

with $U = \psi_{c_2 c_1}^m, V = \phi_{c_1 c_2}^m, W = \tilde{\psi}_{c_1 c_3}^m$ and we will look upon the operators $P_{c_2 c_1}^{mn}, Q_{c_1 c_2}^{mn}, S_{c_1 c_3}^{mn}$ as discretizations of diagonal forms of the operators $U_{c_2 c_1}^{mn}, V_{c_1 c_2}^{mn}, W_{c_1 c_3}^{mn}$.

In order to estimate the number of nodes in the discretizations (3.47), (3.49) required to obtain a given accuracy ϵ in the evaluation of the fields ϕ, ψ we will need the following well known fact (see, for example [7]).

Lemma 3.10. For any integer m , n such that $n \geq 2|m|$, the n -point trapezoidal quadrature rule on the interval $[0, 2\pi]$ integrates the function e^{imx} exactly.

Remark 3.5

By combining the above lemma with Remark 3.1, it is easy to see that the number n of nodes in the discretization G_{ϕ, c_1}^n of the function $G_{\phi, c_1} : [0, 2\pi] \rightarrow C^1$ has to be approximately equal to $2|k|R_1$, and is almost independent of the accuracy ϵ with which the field ϕ is being calculated.

Lemmas 3.5 - 3.7 provide a tool for shifting the origins of asymptotic expansions of radiation fields, and for converting asymptotic representations of the form (3.32) into asymptotic representations of the form (3.38) for a cost proportional to n , where n is the number of nodes in the discretization (3.48) of the interval $[0, 2\pi]$. In the following two sections, this apparatus is used to construct an algorithm for rapid evaluation of integral operators of Section 2.

4. Rapid Evaluation of Integral Operators of Section 2

In this section, we describe an algorithm for rapid evaluation of the field and the normal derivative of the field created on a curve γ by charge and dipole distributions on that same curve. For definitiveness, we will be discussing the evaluation of the field created by a charge distribution. The algorithms evaluating the normal derivative of the field created by a charge distribution, and the field and the normal derivative of the field created by a dipole distribution are quite similar.

4.1. Notation

We will consider the situation depicted in Figure 2. The curve γ is discretized into equispaced nodes x_1, x_2, \dots, x_n , and we will denote the spacing between the adjacent nodes by h . Suppose that for each $i = 1, 2, \dots, n$, a charge a_i of strength σ_i is located at the point x_i . In this section, we describe an algorithm for rapid calculation of approximations $g_i = 1, 2, \dots, n$ to the sums

$$G_\sigma(x_i) = \sum_{\substack{j=1 \\ j \neq i}}^n \sigma_j \phi_{x_j}^k(x_i) \quad (4.1)$$

for $i = 1, 2, \dots, n$. Clearly, this is an order n^2 process (evaluating n fields at n points). However, if we are interested in evaluating (4.1) with a finite accuracy (which is always the case in practical calculations), Theorem 3.1 and Lemmas 3.7, 3.8, 3.9 can be used to speed up the process.

For an integer $m \geq 2$, we will define the points t_1, t_2, \dots, t_{m+1} on the interval $[0, L]$ by the formula $t_i = (i-1)L/m$, subdividing the interval $[0, L]$ into m segments of equal length, and denote the center of the i -th segment by z_i , so that $z_i = t_i + L/(2m)$. For each natural $j = 1, 2, \dots, m$, we will denote by A_j the set of all charges a_i such that $x_i \in \gamma([t_j, t_{j+1}])$, and by D_j the circle of radius $r = L/(2m)$ with the center at $\gamma(z_j)$. We will denote by W_j the union of all A_i such that $\|z_j - z_i\| > 3r$, and by \bar{W}_j the union of all A_i such that $\|z_j - z_i\| \leq 3r$. Obviously, $A_j \subset D_j$ for any $j = 1, 2, \dots, m$. Also, it follows from the triangle inequality that

$$\min_{\substack{x \in A_i \\ y \in A_j}} \|x - y\| \geq r$$

for any i, j such that $A_i \subset W_j$. Finally, we will denote by ϕ_j the field of all charges a_i such that $x_i \in A_j$ and observe that if $x_p \in A_j$ then

$$G_\sigma(x_p) = \sum_{A_i \subset W_j} \phi_i(x_p) + \sum_{x_i \in W_j} \sigma_i \phi_{x_i}^k(x_p). \quad (4.2)$$

4.2. Detailed description of an order $n^{3/2}$ algorithm.

In this subsection, M, N will denote "sufficiently large" integer numbers. The actual choice of the numbers M, N is discussed in the following subsection.

We will evaluate the fields (4.1) in five steps.

Step 1.

Using Lemma 3.8, obtain discretized asymptotic representations $F_{\phi_j, \gamma(z_j)}^N$ of the fields ϕ_j for all $j = 1, 2, \dots, m$.

Step 2

For every pair of natural numbers $i, j \in [1, m]$ such that $A_i \subset W_j$, calculate the representation

$$G_{\psi_{ij}, \gamma(z_j)}^N = S_{\gamma(z_i), \gamma(z_j)}^{M, N} (F_{\phi_j, \gamma(z_j)}^N) \quad (4.3)$$

of the field $\psi_{ij} = \bar{\phi}_{\gamma(z_i), \gamma(z_j)}^M$ and view it as a finite-dimensional approximation to the asymptotic representation of the field ϕ_i on D_j .

Step 3

For each natural $j \in [1, m]$, calculate the sum

$$G_{\psi_j, \gamma(z_j)}^N = \sum_{A_i \subset W_j} G_{\psi_{ij}, \gamma(z_j)}^N \quad (4.4)$$

and view the field $\psi_j = \sum_i \psi_{ij}$ as an approximation to the field $\sum_{A_i \subset W_j} \phi_i$, and $G_{\psi_j, \gamma(z_j)}^N$ as a finite-dimensional approximation to the asymptotic representation of ψ_j on D_j .

Step 4

For each natural $j \in [1, m]$, evaluate

$$\tilde{\psi}_j(x_i) = \tilde{G}_{\psi_j, \gamma(z_j)}(x_i) \quad (4.5)$$

for all i such that $x_i \in \gamma([t_j, t_{j+1}])$ and look upon (4.5) as an approximation to $\psi_j(x_i)$.

Step 5.

For each $j = 1, 2, \dots, m$, evaluate the sum

$$\tilde{\psi}_j(x_i) + \sum_{z_p \in W_j} \sigma_p \phi_{z_p}^k(x_i) \quad (4.6)$$

for all i such that $x_i \in \gamma([t_j, t_{j+1}])$, and view (4.6) as an approximation to $G_\sigma(x_i)$.

4.3. Choice of parameters and CPU time estimate

In the estimates below, a, b, c, d, e are coefficients determined by the computer system, language, particular implementation of the algorithm, etc.

Step 1

Obviously, this step will require order $n \cdot N$ operations (tabulating $F_{\phi_j, \gamma(z_j)}^N$ at N nodes on the interval $[0, 2\pi]$ for each of the nodes x_1, x_2, \dots, x_n). According to the Remark 3.5, $N \sim |k| \cdot L/m$, and the CPU time estimate for this step becomes $a \cdot n \cdot |k| \cdot L/m$.

Step 2

For each of the pairs i, j such that $A_j \subset W_i$, evaluating (4.3) will require order N operations (see (3.54)), and the total number of such pairs is less than m^2 , which results in the CPU time estimate of $b \cdot m^2 \cdot n \sim b \cdot m^2 \cdot |k| \cdot L/m = b \cdot m \cdot |k| \cdot L$ for this step.

Step 3

Obviously, evaluating the sums (4.4) for all $j = 1, 2, \dots, m$ is an order $c \cdot m \cdot N \sim c \cdot m \cdot |k| \cdot L/m = c \cdot |k| \cdot L$ procedure.

Step 4

Evaluating (4.5) for each $i = 1, 2, \dots, n$ is an order N procedure, resulting in the total CPU time estimate for this step of $d \cdot n \cdot N \sim d \cdot n \cdot |k| \cdot L/m$.

Step 5

Evaluating the sum (4.6) for each $i = 1, 2, \dots, n$ is an order n/m procedure, with the resulting CPU time estimate of $e \cdot n^2/m$ for this step.

Summing up the time estimates for the steps 1-5, we obtain the following time estimate for the whole process:

$$T = A \cdot n \cdot |k| \cdot L/m + b \cdot m \cdot |k| \cdot L + c \cdot |k| \cdot L + \frac{e \cdot n^2}{m} \quad (4.7)$$

with $A = a + d$, and we would like to choose m in such a manner that (4.7) would be minimized. Differentiating (4.7) with respect to m , and setting the resulting derivative to zero, we obtain

$$m_{\min} = \sqrt{\frac{A \cdot n \cdot |k| \cdot L + e \cdot n^2}{b \cdot |k| \cdot L}} \quad (4.8)$$

and the corresponding minimum of (4.7) is equal to

$$T_{\min} = 2\sqrt{A \cdot n \cdot |k| \cdot L + e \cdot n^2} \cdot \sqrt{b \cdot |k| \cdot L} + c \cdot k \cdot L. \quad (4.9)$$

If the calculations are performed with a fixed number of nodes per wavelength (which is often a reasonable assumption), n is proportional to $|k|L$, and (4.9) assumes the form

$$T_{\min} \sim (k \cdot L)^{\frac{5}{2}} \quad (4.10)$$

or

$$T_{\min} \sim n^{\frac{5}{2}} \quad (4.11)$$

which for large n is considerably smaller than n^2 .

4.4. Further reduction of the CPU time estimate of the process

The approach of the above subsection can be used recursively by subdividing each of the sets A_i into subsets $\{B_{ij}\}$, $j = 1, 2, \dots, m$ with appropriately chosen m and representing the fields ϕ_i as sums $\phi_i = \sum_j \phi_{ij}$ where ϕ_{ij} is the field created by all charges a_p such that $a_p \in B_{ij}$. A calculation similar to the one in the preceding section shows that such an algorithm will have an asymptotic CPU time estimate of $n^{4/3}$.

In [18], such subdivision process is used recursively to obtain an order n algorithm for numerical evaluation of integral operators of classical potential theory (Laplace's equation). By reproducing the construction of Section VII of [18] almost literally, one can obtain an order $n \log(n)$ algorithm for evaluating (4.1). However, our estimates indicate that for problems of practicable size ($n \leq 20,000$), the improvement in actual computation times obtained by replacing an order $n^{4/3}$ algorithm with an order $n \log n$ algorithm would not be very significant.

5. Implementation and Numerical Results

5.1. Implementation.

In implementing the algorithm of the present paper, we closely followed [16], as far as the formulation of the system of integral equations (2.13), (2.14) and their discretization are concerned.

Equations (2.13), (2.14) were discretized by means of the Nyström algorithm, which calls for replacing the integrals with finite quadrature formulae and the integral equations with systems of linear algebraic equations respectively, (see [3, 16]). Since the kernels of the equations (2.13), (2.14) are logarithmically singular, quadrature formulae suitable for functions with logarithmic singularities have to be employed, and we used a fourth order convergent formula of the type discussed in [19]). Discretization of the equations (2.13), (2.14) by means of the Nyström algorithm leads to large scale, asymptotically well-conditioned systems of linear algebraic equations (see [3]), to which we apply a version of the Generalized Conjugate Residual algorithm (see [11]). Solution of a linear system by means of a Generalized Conjugate Residual type algorithm involves application of the matrix of the system to a sequence of recursively generated vectors, which is accomplished by means of the algorithm of the preceding section (we implemented an order $n^{4/3}$ version of the process). As is well known, in order to converge to a relative accuracy ϵ , a Generalized Conjugate Residual type algorithm requires order $-K \cdot \log(\epsilon)$ iterations, where K is the condition number of the matrix of the system being solved. This results in the total CPU time estimate of the order $n^{4/3} \cdot \log(\epsilon)$ for the solution of the equations (2.13), (2.14), since the application of the Nyström algorithm to these equations leads to well-conditioned linear systems (see [3]). After the system of equations (2.13), (2.14) is solved, evaluating the scattered field at any point $x \notin \Gamma$ involves two numerical integrations over Γ (evaluating the field of the charge distribution and that of the dipole distribution), which is an order n procedure. This brings us to the estimate

$$T_{SOLVE} = -a \cdot \log(\epsilon) \cdot n^{4/3} + b \cdot n \cdot m$$

for obtaining the solution of the problem (2.6) - (2.10), where m is the number of points in R^2 where the solution has to be calculated.

5.2. Numerical Results.

We have applied the algorithm of the present paper to several acoustic scattering problems in two dimensions. Three examples are presented below.

a. Scattering from a circular inclusion. In this case, the problem possesses an analytical solution, providing a convenient way to verify the accuracy of the algorithm. In our example, the radius of the scatterer was 100 ft., and its center was at the origin. The densities inside and outside the scatterer were 1.2 and 1.0 respectively, and the speeds of sound were 20 and 18 respectively. The ambient field was generated by a cylindrically symmetric source located at the point (100,100), and the scattered field obtained by the algorithm of the present paper was compared to the analytical result at 40 equispaced points located on the circle of radius 110 with the center at the origin (see Figure 3). The calculation was performed with the angular frequency of the source varying from 10 to 320, and in all cases the boundary of the scatterer was discretized at 10 nodes per wavelength, which usually results in three to four digit accuracy. In Table 1, ω is the angular frequency, n is the number of nodes on the boundary of the scatterer (at 10 nodes per wavelength), ϵ is the resulting mean square error at 40 receivers, and m is the number of iterations the Generalized Conjugate Residual algorithm took to converge to 4-digit accuracy. Following observations can be made from Table 1.

- 1.. The accuracy of the solution obtained with a given number of nodes per wavelength is virtually independent of the frequency.
- 2.. The number of iterations required by the GCRA to obtain a given accuracy is almost independent of the frequency.
- 3.. The actual CPU times required by the algorithm grow somewhat erratically as a function of the number of nodes on the boundary, but seem to be in agreement with the theoretical CPU time estimate of the algorithm.

b. Scattering from a lense-shaped object. In the situation depicted in Figure 4, the speed of sound in the containing space is 30 ft/sec, and the speed of sound inside the scatterer is 20 ft/sec. Both upper and lower surfaces of the scatterer are circular with the radius of curvature 1 ft, so that the focal distance of the resulting lense is 2 feet. Three cylindrically symmetric sources are located at the points I, II, and III respectively have the frequency $F = 7162$ Hz, resulting in the wavelengths 0.0041888 outside the scatterer, and 0.0027925 ft inside the scatterer. The amplitude of the field generated by this configuration on the screen is depicted in Figure 5. Clearly, the laws of geometrical optics should be applicable to this situation with a reasonable degree of accuracy, since the lense is about 240 exterior (or 360 interior) wavelengths in diameter, and a careful examination of the Figure 5 shows this to be the case. In order to produce the Figure 5, the surface of the lense in Figure 4 was discretized into 6134 nodes (with 10 nodes per wavelength), resulting in a system of linear algebraic equations of dimension 12268. The solution of this problem took 3358 seconds of CPU time on a VAX-8600.

c. Radiation from a worm-shaped object. In this numerical experiment, a source of angular frequency $\omega = 640$ was located inside an inclusion of a complicated shape in a fluid two-dimensional space (see Figure 6). The densities inside and outside the inclusion were 1.25 and 1.0 respectively, and the speeds of sound were 20 and 25 respectively. The boundary of the inclusion was discretized at 10 nodes per wavelength, resulting in the total number of 5078 nodes on the boundary of the inclusion. The calculation took 3867 seconds of CPU time, and the amplitude of the field as measured by receivers in Figure is displayed in Figure 7.

6. Generalizations and Conclusions

6.1. Other boundary conditions.

In the present paper, we have been solving the problem (2.6) - (2.10), corresponding to the scattering of sound from a fluid inclusion in a fluid space. Obviously, the two classical acoustical scattering problems (Dirichlet and Neumann problems for the Helmholtz equation, corresponding to scattering from a cavity and from a rigid inclusion respectively) can be solved in a similar manner. Generally, whenever a scattering problem is reduced to a set of second kind integral equations on the boundary of the scatterer by means of some form of Stoke's theorem, the algorithm of the Section IV provides a tool for solving these equations in order $n^{4/3}$ operations.

6.2. Multiple scatterers.

The case of multiple scatterers does not differ substantially from the case of a single scatterer. The scattered field inside each scatterer is represented in a manner precisely similar to the one used to solve a single scatterer problem. The scattered field outside the scatterers is represented by the sum of the fields of single and double layer potentials on the surfaces of all scatterers. The whole procedure is completely straightforward.

6.3. Three-dimensional version of the theory.

Three-dimensional equivalents of the expansions (3.1), (3.2) are well known (see, for example, [15]), and those of the mappings $U_{c_2c_1}^n, V_{c_1c_2}^n, W_{c_1c_3}^n$ are fairly easy to define. However, in two dimensions, the translation operators $U_{c_2c_1}^n, V_{c_1c_2}^n, W_{c_1c_3}^n$ are convolutions, and the diagonal form of the latter is well known from the standard theory of the Fourier integral (see Theorem 3.1). In three dimensions, no such ready-made analytical apparatus is available, and the proof of an equivalent of the Theorem 3.1 is considerably more involved. Fortunately, diagonal forms of the three-dimensional analogues of the operators $U_{c_2c_1}^n, V_{c_1c_2}^n, W_{c_1c_3}^n$ can be obtained by a generalization of the technique used in [20] to derive Graf's addition theorem, permitting fast solvers for scattering problems in three dimensions to be constructed. Details of this generalization will be reported at a later date.

6.4. Conclusions

One of principal difficulties arising in the solution of large-scale scattering problems by means of integral equations is the fact that the Green's function for the Helmholtz equation decays slowly. As a result, the kernels of the obtained integral equations are not sparse, and their discretization leads to dense large scale systems of linear algebraic equations. Solution of such systems is an extremely expensive process (see [9, 11]), which limits the usefulness of the whole approach.

In Section 4 of the present paper, we construct an algorithm for rapid application of the matrices resulting from discretization of integral equations of scattering theory to arbitrary vectors. The asymptotic CPU time estimate of the algorithm is $n^{4/3}$, and can be reduced to $n \log(n)$ where n is the number of nodes in the discretization of the boundary of the scatterer. By combining the approach of the Section 4 with a Generalized Conjugate Residual type process, we obtain an order $n^{1/3} \log(\epsilon)$ algorithm for the solution of the integral equations (2.13), (2.14) of acoustic scattering theory. This results in acceptable computation times even for large-scale scattering problems, (see preceding section), as long as the solution has to be evaluated at a limited number of points.

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Table 1

ω	n	ϵ	m	$T(sec.)$
10	128	$0.766E-3$	5	36
20	257	$0.594E-3$	6	83
40	514	$0.629E-3$	6	179
80	1028	$0.620E-3$	6	612
160	2056	$0.708E-3$	6	1696
320	4112	$0.727E-3$	6	4328

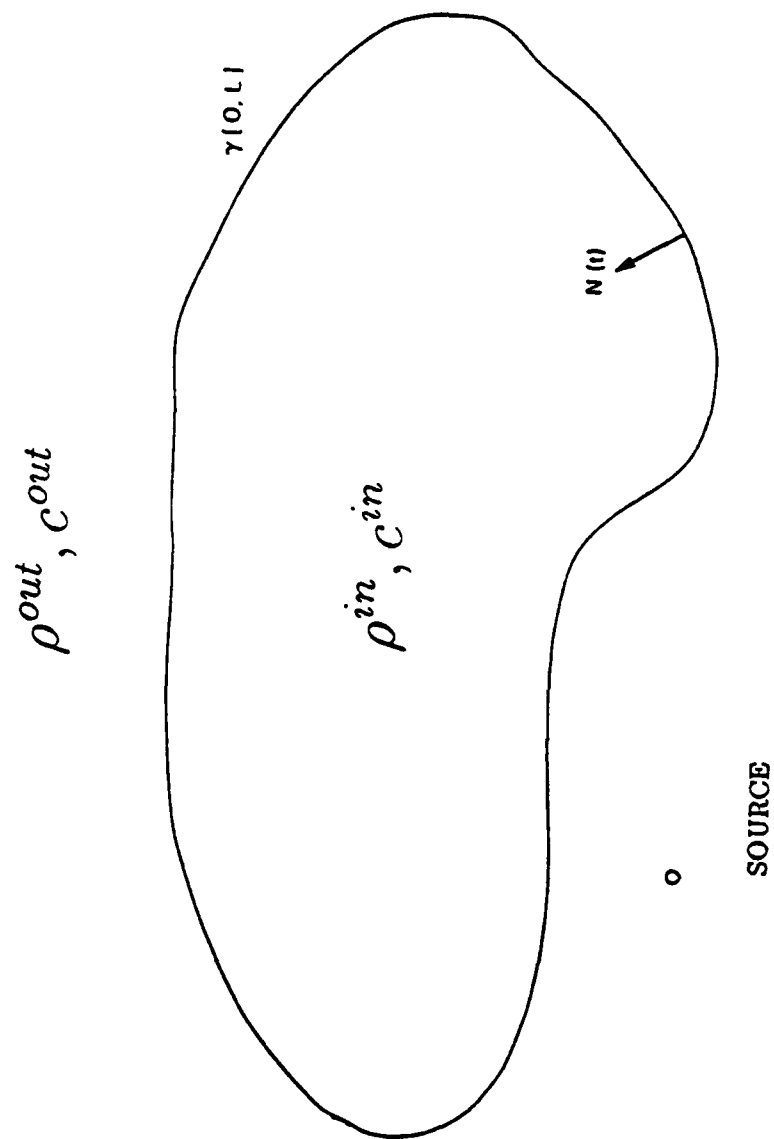


Figure 1

A Fluid Scatterer Imbedded in a Two-dimensional Fluid Space

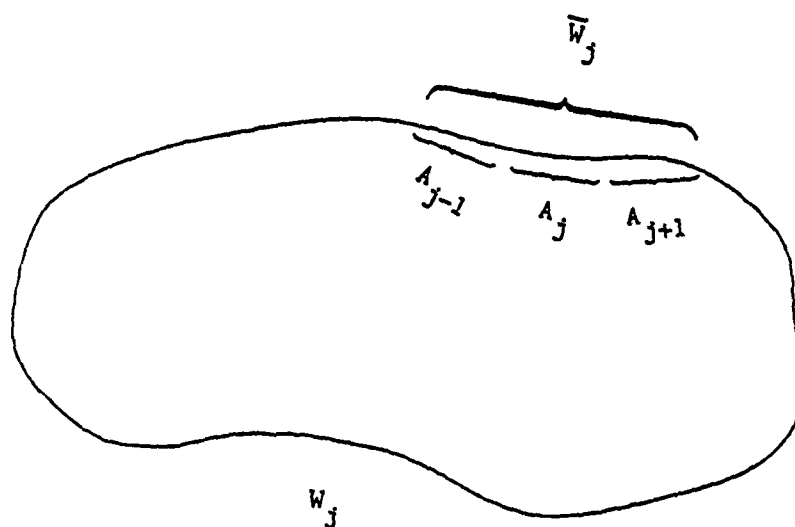


Figure 2

Definition of the Sets A_j , \bar{W}_j , W_j

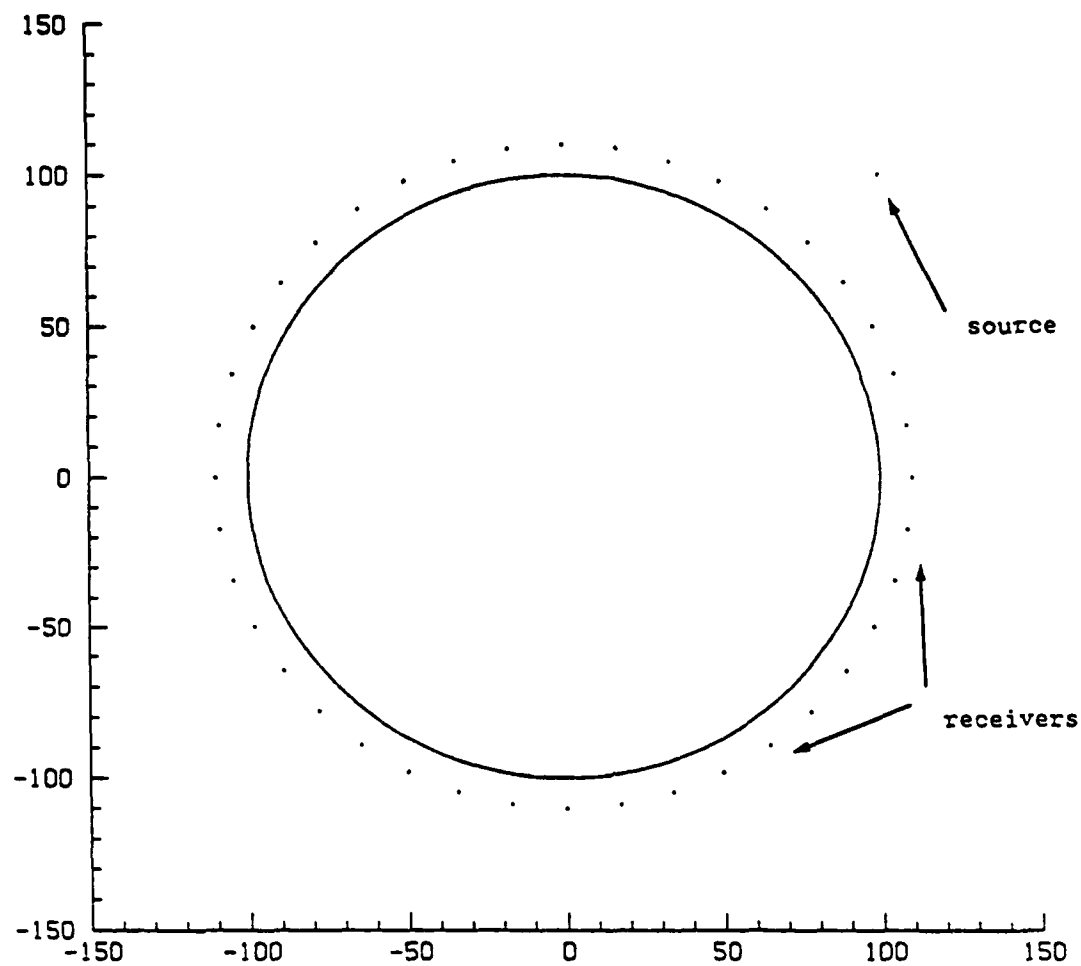


Figure 3

Acoustic Scattering From a Circular Inclusion

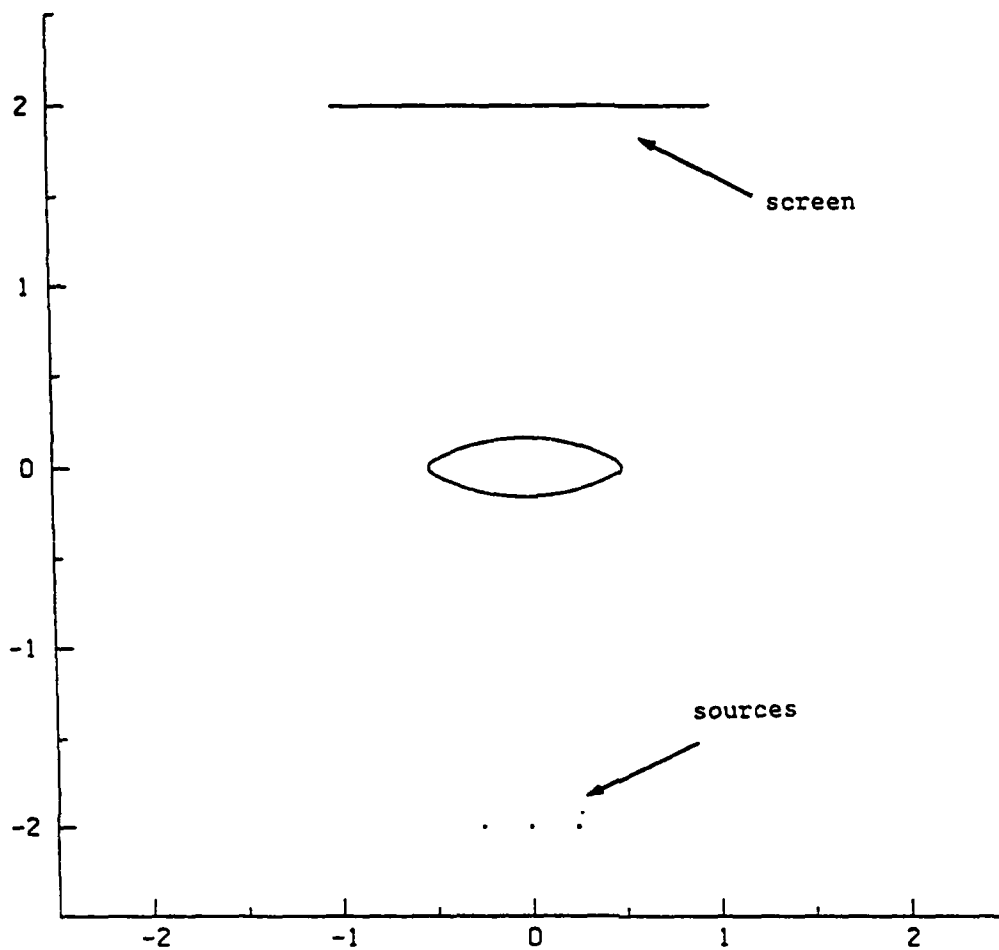


Figure 4

Acoustic Scattering From a Lense-Shaped Inclusion

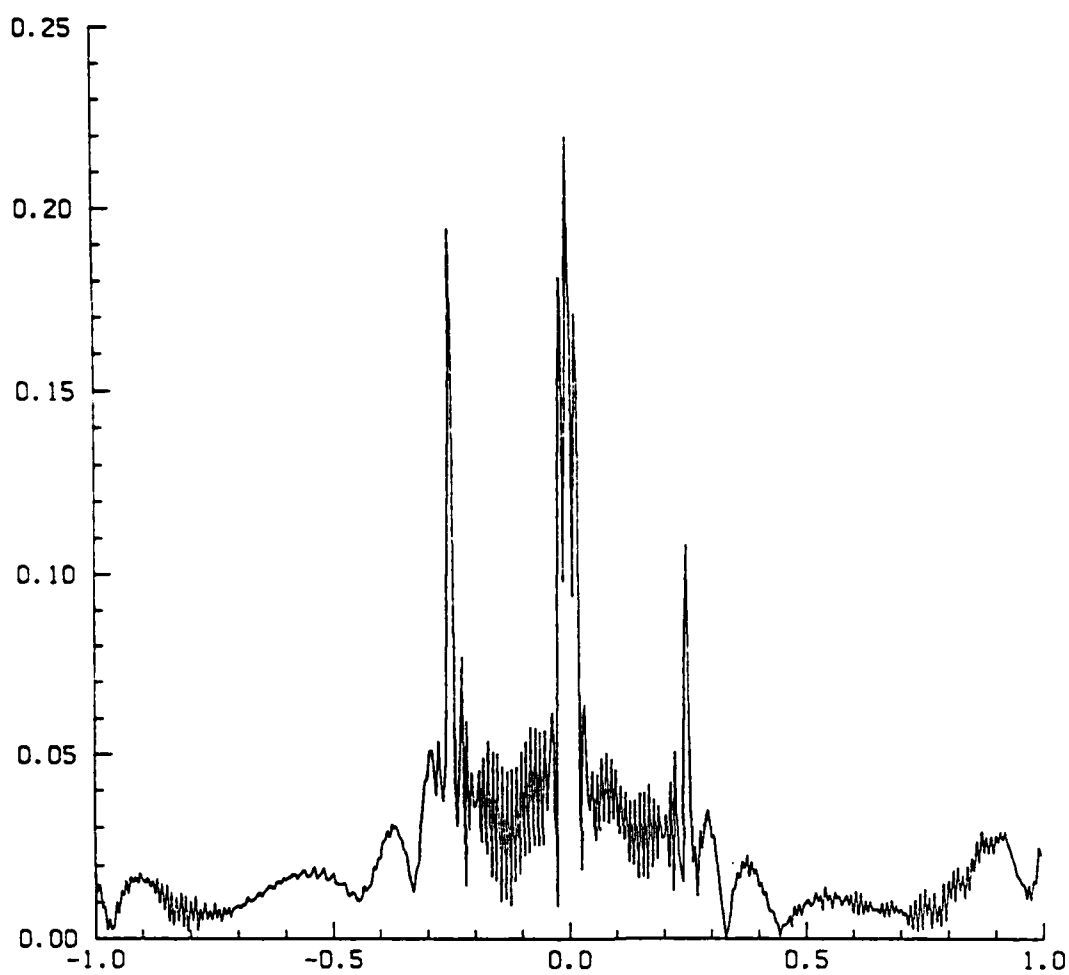


Figure 5

Amplitude of the Field Measured Along the Screen in Figure 4

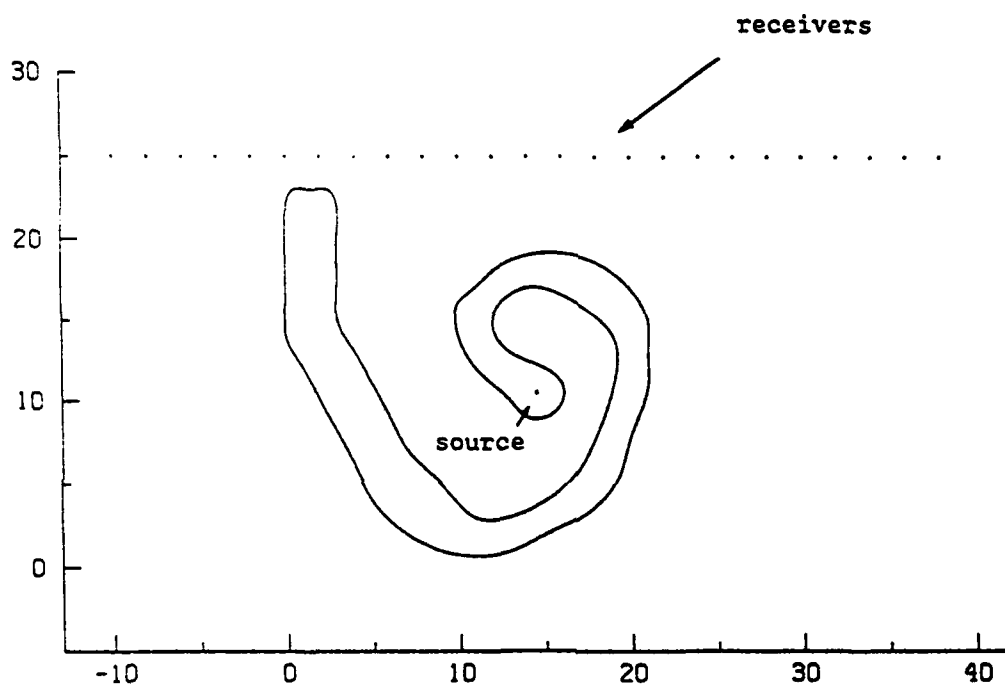


Figure 6

Acoustic Radiation From a Worm-Shaped Inclusion

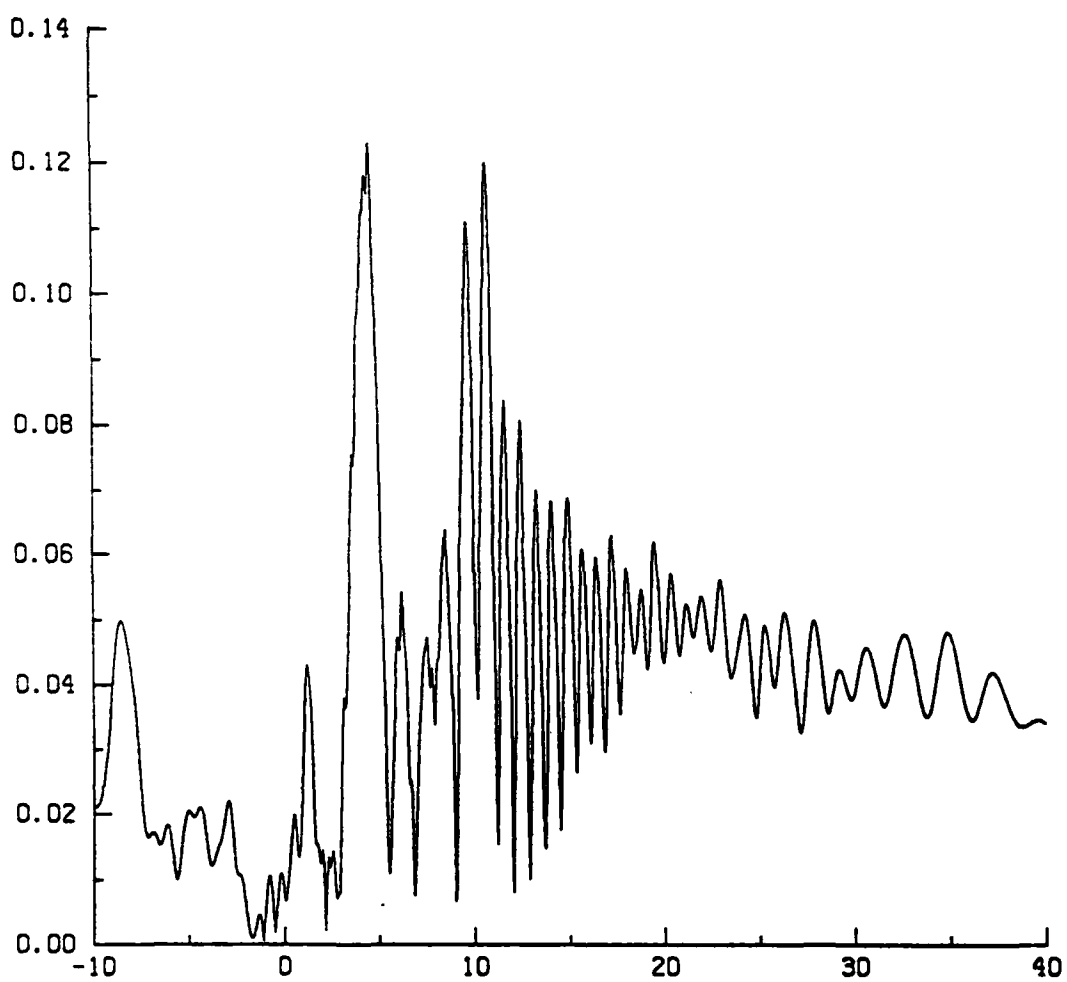


Figure 7

Amplitude of the Field Measured by Receivers in Figure 6

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